# Lawrence Berkeley National Laboratory

LBL Publications

Title Minimal non-abelian supersymmetric Twin Higgs

Permalink

https://escholarship.org/uc/item/2f51r78k

Journal

Journal of High Energy Physics, 2017(10)

ISSN

1126-6708

Authors

Badziak, Marcin Harigaya, Keisuke

Publication Date

2017-10-01

DOI

10.1007/jhep10(2017)109

Peer reviewed

Published for SISSA by 🖄 Springer

RECEIVED: August 13, 2017 REVISED: September 21, 2017 ACCEPTED: September 29, 2017 PUBLISHED: October 17, 2017

# Minimal non-abelian supersymmetric Twin Higgs

Marcin Badziak<sup>*a,b,c*</sup> and Keisuke Harigaya<sup>*b,c*</sup>

- <sup>a</sup>Institute of Theoretical Physics, Faculty of Physics, University of Warsaw, ul. Pasteura 5, PL-02-093 Warsaw, Poland
- <sup>b</sup>Department of Physics, University of California, Berkeley, California 94720, U.S.A.
- <sup>c</sup> Theoretical Physics Group, Lawrence Berkeley National Laboratory, Berkeley, California 94720, U.S.A.

*E-mail:* mbadziak@berkeley.edu, keisukeharigaya@berkeley.edu

ABSTRACT: We propose a minimal supersymmetric Twin Higgs model that can accommodate tuning of the electroweak scale for heavy stops better than 10% with high mediation scales of supersymmetry breaking. A crucial ingredient of this model is a new  $SU(2)_X$ gauge symmetry which provides a *D*-term potential that generates a large SU(4) invariant coupling for the Higgs sector and only small set of particles charged under  $SU(2)_X$ , which allows the model to be perturbative around the Planck scale. The new gauge interaction drives the top yukawa coupling small at higher energy scales, which also reduces the tuning.

KEYWORDS: Beyond Standard Model, Supersymmetric Standard Model

ARXIV EPRINT: 1707.09071



#### Contents

Testus duration

| 1            | Introduction   | 1 |  |  |  |  |
|--------------|--|---|--|--|--|--|
| <b>2</b>     | A SUSY $D$ -term Twin Higgs with an SU(2) gauge symmetry | 2 |  |  |  |  |
|              | 2.1 $SU(2)_X$ symmetry breaking                          | 4 |  |  |  |  |
|              | 2.2 SU(4) invariant quartic coupling and $\mu$ terms     | 5 |  |  |  |  |
|              | 2.3 Masses of matter particles                           | 7 |  |  |  |  |
| 3            | Fine-tuning of the electroweak scale                     |   |  |  |  |  |
| 4            | l Summary  |   |  |  |  |  |
| $\mathbf{A}$ | A Electroweak precision measurements                     |   |  |  |  |  |

#### 1 Introduction

Supersymmetry (SUSY) provides one of the most promising solutions to the hierarchy problem of the Standard Model (SM) [1–4]. However, the lack of finding of SUSY partners casts serious doubts on whether SUSY can still naturally explain the electroweak (EW) scale. Fine-tuning of the EW scale in minimal SUSY models implied by the LHC searches was recently quantified in refs. [5, 6], which demonstrated that the current limits on stop and gluino masses exclude regions with fine-tuning better than 10%, even if a very low mediation scale of the SUSY breaking of 100 TeV is assumed.<sup>1</sup> The fine-tuning quickly gets worse for larger mediation scales due to longer RG running of the soft Higgs mass. This is indication of the little hierarchy problem.

A possible remedy to the little hierarchy problem is offered by Twin Higgs mechanism [11-15]. In the scenario, the Higgs is a pseudo-Nambu-Goldstone boson of a global SU(4) symmetry emerging from  $\mathbb{Z}_2$  symmetry exchanging the SM with its mirror (or twin) copy. We refer to [16-23] for composite Twin Higgs models, and [24-34] for cosmological aspects of Twin Higgs scenario.

Early realisations of SUSY UV completion of Twin Higgs scenario [14, 15], which generate an SU(4) invariant quartic term with an F-term potential of a heavy singlet superfield, are not able to significantly reduce fine-tuning as compared to non-Twin SUSY models [35-37]. It was only very recently that SUSY Twin Higgs models were proposed in which tuning at the level of 10% is possible by introducing either hard  $\mathbb{Z}_2$  symmetry breaking in the F-term model [36] or a new  $U(1)_X$  gauge symmetry whose D-term potential

<sup>&</sup>lt;sup>1</sup>The fine-tuning may be improved if the higgsino mass is not tied to the Higgs mass squared, see e.g. refs. [7-10]. In such a case higgsino could be heavier leading to compressed spectra for which the lower bounds on stops, and gluino are much weaker.

provides a large SU(4) invariant quartic term [37]. It should be, however, emphasised that the tuning at the level of 10% can be obtained in these models only for a low mediation scale or a low Landau pole scale. In the *F*-term model of ref. [36] a fine-tuning penalty for a larger mediation scale and hence a longer RG running is severe because the large SU(4) invariant coupling induces growth of the top yukawa coupling at higher energy scales. In the *D*-term model the RG effect of the gauge coupling  $g_X$  of the new interaction is to reduce the top yukawa coupling, and the effect of a higher mediation scale is not as severe as the one for the *F*-term model. However, the RG running of the U(1)<sub>X</sub> gauge coupling is fast and hence the Landau pole scale of  $g_X$  is as low as  $10^5-10^6$  GeV for values of  $g_X$  that are large enough to guarantee approximate SU(4) symmetry of the Higgs potential. While such a low mediation scale or a low Landau pole scale is in principle possible, it strongly limits possible schemes of the mediation of the SUSY breaking and UV completions above the Landau pole scale.

In the present work, we point out that the Landau pole scale and the mediation scale of the *D*-term model can be much higher if the SU(4) invariant term is generated by a *D*-term potential of a new non-abelian gauge symmetry. We construct a consistent model with  $SU(2)_X$  gauge symmetry with small number of flavors charged under this symmetry. The new gauge interaction drives the top yukawa coupling small at higher energy scales, which also helps obtain the EW scale more naturally. As a result, the tuning of the EW scale for 2 TeV stops and gluino can be at the level of 5–10% for mediation scales as high as  $10^9-10^{13}$  GeV. One can keep perturbativity up to around the Planck scale with tuning better than 5% (for low mediation scales). The model allows for moderate tuning better than few percent with the mediation scale around the Planck scale. If the gluino mass is a Dirac one, the tuning may be as good as 10%, which realizes a natural SUSY with a gravity mediation.

## 2 A SUSY *D*-term Twin Higgs with an SU(2) gauge symmetry

In this section we present a SUSY *D*-term Twin Higgs model [37] where the *D*-term potential of a new  $SU(2)_X$  gauge symmetry generates the SU(4) invariant quartic coupling. We assume a  $\mathbb{Z}_2$  symmetry exchanging the SM with its mirror copy, and denote mirror objects with superscripts '.

The matter content of the model is shown in table 1. In addition to the  $SU(3)_c \times SU(2)_L \times U(1)_Y$  gauge symmetry and its mirror counterpart, we introduce an  $SU(2)_X$  gauge symmetry which is neutral under the  $\mathbb{Z}_2$  symmetry. We embed an up-type Higgs  $H_u$  into a bi-fundamental of  $SU(2)_L \times SU(2)_X$ ,  $\mathcal{H}$ , and its mirror partner  $H'_u$  into that of  $SU(2)'_L \times SU(2)_X$ ,  $\mathcal{H}'$ . As we will see later, the *D*-term potential of  $SU(2)_X$  is responsible for the SU(4) invariant quartic coupling of  $H_u$  and  $H'_u$ . The  $SU(2)_X$  symmetry is broken by the vacuum expectation value (VEV) of a pair of  $SU(2)_X$  fundamental *S* and  $\bar{S}$ . Except for *S* and  $\bar{S}$  all matter fields have their mirror partner.

The right-handed top quark is embedded into  $\bar{Q}_R$  and allow for a large enough top yukawa coupling through the superpotential term  $\mathcal{H}\bar{Q}_RQ_3$ , where  $Q_3$  is the third generation quark doublet.  $\bar{E}$  is necessary in order to cancel the U(1)<sub>Y</sub>-SU(2)<sup>2</sup><sub>X</sub> anomaly. The VEV of

|                    | $\mathrm{SU}(2)_X$ | $\mathrm{SU}(2)_L$ | ${ m SU}(2)'_L$ | $U(1)_Y$ | $\mathrm{U}(1)_Y'$ | $SU(3)_c$      | $SU(3)_c'$     |
|--------------------|--------------------|--------------------|-----------------|----------|--------------------|----------------|----------------|
| $\mathcal{H}$      | 2                  | 2                  |                 | 1/2      |                    |                |                |
| $\mathcal{H}'$     | 2                  |                    | 2               |          | 1/2                |                |                |
| $\bar{Q}_R$        | 2                  |                    |                 | -2/3     |                    | $\overline{3}$ |                |
| $\bar{Q}'_R$       | 2                  |                    |                 |          | -2/3               |                | $\overline{3}$ |
| S                  | 2                  |                    |                 |          |                    |                |                |
| $\bar{S}$          | 2                  |                    |                 |          |                    |                |                |
| Ē                  | 2                  |                    |                 | 1        |                    |                |                |
| $\bar{E}'$         | 2                  |                    |                 |          | 1                  |                |                |
|                    |                    |                    |                 | 2/3      |                    | 3              |                |
| U'                 |                    |                    |                 |          | 2/3                |                | 3              |
| $E_{1,2}$          |                    |                    |                 | -1       |                    |                |                |
| $E'_{1,2}$         |                    |                    |                 |          | -1                 |                |                |
| $\phi_u$           |                    | 2                  |                 | 1/2      |                    |                |                |
| $\phi'_u$          |                    |                    | 2               |          |                    |                |                |
| $\phi_{d1,2,3}$    |                    | 2                  |                 | -1/2     |                    |                |                |
| $\phi'_{d1,2,3}$   |                    |                    | 2               |          | -1/2               |                |                |
| $Q_{1,2,3}$        |                    | 2                  |                 | 1/6      |                    | 3              |                |
| $\bar{u}_{1,2}$    |                    |                    |                 | -2/3     |                    | $\overline{3}$ |                |
| $\bar{e}_{1,2,3}$  |                    |                    |                 | 1        |                    |                |                |
| $\bar{d}_{1,2,3}$  |                    |                    |                 | 1/3      |                    | $\overline{3}$ |                |
| $L_{1,2,3}$        |                    | 2                  |                 | -1/2     |                    |                |                |
| $Q'_{1,2,3}$       |                    |                    | 2               |          | 1/6                |                | 3              |
| $\bar{u}_{1,2}'$   |                    |                    |                 |          | -2/3               |                | $\overline{3}$ |
| $\bar{e}'_{1,2,3}$ |                    |                    |                 |          | 1                  |                |                |
| $\bar{d}'_{1,2,3}$ |                    |                    |                 |          | 1/3                |                | $\overline{3}$ |
| $L'_{1,2,3}$       |                    |                    | 2               |          | -1/2               |                |                |

JHEP10(2017)109

 $\label{eq:table 1. The matter content of the model.}$ 

 $\phi_u$  is responsible for the masses of the up and charm quarks.  $Q_{1,2,3}$ ,  $\bar{u}_{1,2,3}$ ,  $\bar{e}_{1,2,3}$ ,  $\bar{d}_{1,2,3}$  and  $L_{1,2,3}$  are usual MSSM fields. To cancel the gauge anomaly of  $SU(3)_c^2 - U(1)_Y$  and  $U(1)_Y^3$  originating from the extra up-type right handed quark in  $\bar{Q}_R$  and two extra right-handed leptons in  $\bar{E}$ , we introduce U and  $E_{1,2}$ . There are three up-type Higgses in  $\mathcal{H}$  and  $\phi_u$ , so we need to introduce three down-type Higgsses  $\phi_{d1,2,3}$ . Their VEVs are responsible for the masses of down-type quarks and charged leptons.

# 2.1 $SU(2)_X$ symmetry breaking

We introduce a singlet chiral field Z and the superpotential coupling

$$W = \kappa Z (S\bar{S} - M^2). \tag{2.1}$$

We assume that the soft masses of S and  $\overline{S}$  are the same,

$$V_{\text{soft}} = m_S^2 (|S|^2 + |\bar{S}|^2).$$
(2.2)

Otherwise, the magnitude of the VEVs of S and  $\overline{S}$  are different from each other, and give large soft masses to the Higgs doublets through the *D*-term potential. The VEVs of S and  $\overline{S}$  are given by

$$\langle S \rangle = \begin{pmatrix} 0 \\ v_S \end{pmatrix}, \ \langle \bar{S} \rangle = \begin{pmatrix} v_S \\ 0 \end{pmatrix}, \ v_S = \sqrt{M^2 - m_S^2 / \kappa^2}.$$
 (2.3)

The constraint on the  $T(\rho)$  parameter requires that  $v_S \gtrsim 2.9 \text{ TeV}$  in the limit of large tan  $\beta$ and neglecting the effect of mixing between the SM and the mirror Higgses, see appendix A for a derivation of this constraint and more precise formula. The masses of the  $\text{SU}(2)_X$ gauge bosons are given by

$$m_X^2 = g_X^2 v_S^2. (2.4)$$

After integrating out massive particles with a mass as large as  $v_S$ , the potential of  $\mathcal{H}$ and  $\mathcal{H}'$  is given by

$$\frac{1}{8}g_X^2 \sum_{i=1,2,3} \left( \mathcal{H}^{\dagger} \sigma^i \mathcal{H} + \mathcal{H}'^{\dagger} \sigma^i \mathcal{H}' \right)^2 \left( 1 - \epsilon^2 \right), \qquad (2.5)$$

$$\epsilon^2 = \frac{m_X^2}{2m_S^2 + m_X^2}.$$
 (2.6)

In the SUSY limit,  $m_S^2 = 0$ , the *D*-term potential vanishes. In terms of the model parameters  $M, m_S, \kappa, g_X, \epsilon^2$  is given by

$$\epsilon^2 = \frac{g_X^2 (m_S^2 - \kappa^2 M^2)}{g_X^2 (m_S^2 - \kappa^2 M^2) - 2\kappa^2 m_S^2} \,. \tag{2.7}$$

In the limit where  $\kappa \ll g_X$ ,  $\epsilon^2 = 1$  and hence the *D*-term potential decouples. In order to obtain a large *D*-term potential, it is preferable that  $\kappa$  is as large  $g_X$ .

To estimate the maximal possible value of  $\kappa$ , we solve the renormalization group equation of  $g_X$  and  $\kappa$ ,

$$\frac{\mathrm{d}}{\mathrm{dln}\mu}g_X = \frac{g_X^3}{16\pi^2} \frac{1 + \frac{21}{16\pi^2}g_X^2 - \frac{1}{8\pi^2}\kappa^2}{1 - \frac{g_X^2}{4\pi^2}},\tag{2.8}$$

$$\frac{\mathrm{d}}{\mathrm{dln}\mu}\kappa = \frac{\kappa}{16\pi^2}(4\kappa^2 - 3g_X^2),\tag{2.9}$$



**Figure 1**. Running of  $g_X$  and  $\kappa$  from a high energy scale  $M_*$  down to lower energy scales.

from a high energy scale  $M_*$  towards low energy scales, with a boundary condition at  $M_*$ of  $g_X = \kappa \simeq 2\pi$ .  $M_*$  can be identified with the Landau pole scale. The running of  $g_X$  and  $\kappa$  is shown in figure 1, which shows that  $\kappa \simeq g_X$  much below  $M_*$ . We obtain the same conclusion as long as  $\kappa(M_*) \gtrsim 1$ . For  $\kappa \simeq g_X$ ,  $\epsilon^2$  is

$$\epsilon^2 \simeq \frac{g_X^2 M^2 - m_S^2}{g_X^2 M^2 + m_S^2}.$$
(2.10)

We may obtain a sufficiently small  $\epsilon^2$ , say  $\epsilon^2 \lesssim 0.2$ , for  $m_S^2 \gtrsim 0.6 g_X^2 M^2$ .

Notice also that for  $\epsilon^2 < 1$  there is a threshold correction to the soft Higgs mass which is proportional to a new gauge bosons mass squared:

$$\left(\delta m_{H_u}^2\right)_X = 3 \frac{g_X^2}{64\pi^2} m_X^2 \ln\left(\epsilon^{-2}\right) \,, \tag{2.11}$$

which may be a source of tuning of the EW scale. The same threshold correction is present also for the right-handed stop soft mass squared  $m_{U_2}^2$ .

#### 2.2 SU(4) invariant quartic coupling and $\mu$ terms

We give masses to  $\mathcal{H} = (H_1, H_2)^T$  and  $\phi_u$  by pairing them with  $\phi_{d,1,2,3}$  through the superpotential terms,

$$W = \lambda_1 \phi_{d,1} \mathcal{H}S + \lambda_2 \phi_{d,2} \mathcal{H}\bar{S} + m \phi_u \phi_{d,3} \,. \tag{2.12}$$

The pairs  $(H_1, \phi_{d,1})$ ,  $(H_2, \phi_{d,2})$  and  $(\phi_u, \phi_{d,3})$  obtain masses of  $\lambda_1 v_S$ ,  $\lambda_2 v_S$  and m, respectively. We assume that  $\lambda_2 v_S$ ,  $m \gtrsim 1$  TeV and neglect  $(H_2, \phi_{d,2})$  and  $(\phi_u, \phi_{d,3})$  for the dynamics of the electroweak symmetry breaking. We identify  $H_1$  and  $\phi_{d,1}$  with  $H_u$  and  $H_d$  in the Higgs sector of the standard SUSY model. The  $\mu$  parameter is given by  $\mu = \lambda_1 v_S$ . The SU(4) invariant quartic coupling of  $(H_u, H'_u)$  is given by

$$V = \frac{g_X^2}{8} (1 - \epsilon^2) (|H_u|^2 + |H'_u|^2)^2.$$
(2.13)

As we will see, the VEV of  $\phi_u$  is responsible for the masses of the up and charm quarks, and the neutrinos. To give a VEV to  $\phi_u$ , we introduce a coupling

$$W = \delta m \phi_u \phi_{d,1}. \tag{2.14}$$

Through the F term potential of  $\phi_{d,1}$ ,  $\phi_u$  obtains a tadpole term after  $H_u$  obtains its VEV, which induces a non-zero VEV of  $\phi_u$ .

Through the coupling  $\lambda_2(>\lambda_1)$ ,  $m_{H_u}^2$  receive a quantum correction from  $m_S^2$ ,

$$\Delta m_{H_u}^2 \simeq -\frac{\lambda_2^2}{8\pi^2} m_S^2 L = -(600 \text{ GeV})^2 \left(\frac{\lambda_2}{0.3}\right)^2 \frac{m_S^2}{(6 \text{ TeV})^2} \frac{L}{\ln 10^4}, \qquad (2.15)$$

where L denotes a log-enhancement through an RGE. As long as  $\lambda_2 \lesssim 0.4$ , this contribution is always smaller than that from stops and/or the threshold correction from X, and hence we neglect it. Note, however, that even larger values of  $\lambda_2$  may be possible without introducing tuning if the mediation scale of SUSY breaking is relatively low and/or  $m_S^2$ runs to smaller values at higher energies.

Note that the  $Z_2$  symmetry  $S \leftrightarrow \overline{S}$  is explicitly broken by the above superpotential couplings. Even if we assume the  $Z_2$  symmetry of the soft masses of S and  $\overline{S}$ , we expect a quantum correction to a mass difference of them,

$$\Delta m_S^2 \equiv m_S^2 - m_{\bar{S}}^2 \simeq \frac{\lambda_2^2}{8\pi^2} m_S^2 L.$$
 (2.16)

This leads a asymmetric VEV of S and  $\bar{S}$ , which give  $m_{H_u}^2$  through the D-term potential,

$$m_{H_u^2} \simeq -\frac{\epsilon^2}{2} \Delta m_S^2, \qquad (2.17)$$

which is always smaller than the direct one-loop quantum correction in eq. (2.15).

It is also possible to maintain the  $Z_2$  symmetry. Instead of the coupling in eqs. (2.12) and (2.14), we introduce

$$W = \mathcal{H}(\lambda_1 \phi_{d,1} + \lambda_3 \phi_{d,3})(S + \bar{S}) + \lambda_2 \mathcal{H}\phi_{d,2}(S - \bar{S}) + \phi_u(m_1 \phi_{d,1} + m_3 \phi_{d,3}).$$
(2.18)

Here we have assumed that  $\phi_{d,2}$  is odd under the  $Z_2$  symmetry. After S and  $\overline{S}$  obtain their VEVs, the mass terms become

$$W = v_S(\lambda_1\phi_{d,1} + \lambda_3\phi_{d,3})(H_1 - H_2) + \lambda_2 v_S\phi_{d,2}(H_1 + H_2) + \phi_u(m_1\phi_{d,1} + m_3\phi_{d,3}).$$
(2.19)

We assume that  $\lambda_2 v_S$ ,  $m_i \gtrsim 1$  TeV. Then  $(H_1 + H_2)/\sqrt{2}$  and  $\phi_u$  obtain a large mass paired with  $\phi_{d,2}$  and a linear combination of  $\phi_{d,1}$  and  $\phi_{d,3}$ , respectively, and are irrelevant for the dynamics of the electroweak symmetry breaking.  $H_u \equiv (H_1 - H_2)/\sqrt{2}$  obtains a mass of  $O(\lambda_{2,3}v_S)$  paired with another linear combination of  $\phi_{d,1}$  and  $\phi_{d,3}$  which we call  $H_d$ .

#### 2.3 Masses of matter particles

We first consider a case where the  $Z_2$  symmetry  $S \leftrightarrow \overline{S}$  is explicitly broken. A large enough top yukawa coupling is obtained by the superpotential

$$W = y_t \mathcal{H}\bar{Q}_R Q_3 \to y_t \left( H_2 \bar{Q}_{R,1} - H_1 \bar{Q}_{R,2} \right) Q_3, \tag{2.20}$$

where  $\bar{Q}_R = (\bar{Q}_{R,1}, \bar{Q}_{R,2})^T$ . We give a large mass to  $\bar{Q}_{R,1}$  by introducing a coupling

$$W = y\bar{Q}_R US,\tag{2.21}$$

and identify  $\bar{Q}_{R,2}$  with a right-handed top quark  $\bar{u}_3$ .

The yukawa couplings of the up and charm quarks originates from the couplings with  $\phi_u$ ,

$$W = y_{u,ij}\phi_u Q_i \bar{u_j}.$$
(2.22)

The left-handed neutrino masses are obtained in a similar manner once right-handed neutrinos are introduced. The yukawa couplings of the down-type quarks and the charged leptons is given by couplings with  $\phi_{d,i}$ ,

$$W = y_{d,ijk}\phi_{d,i}Q_jd_k + y_{e,ijk}\phi_{d,i}L_j\bar{e}_k.$$
(2.23)

The extra  $SU(2)_X$  charged particle  $\overline{E}$  obtains its mass paired with  $E_{1,2}$  through the  $SU(2)_X$  symmetry breaking,

$$W = \bar{E}(y_{E,1}E_1 + y_{E,2}E_2)S + \bar{E}(\bar{y}_{E,1}E_1 + \bar{y}_{E,2}E_2)\bar{S}.$$
(2.24)

Next we consider a case where the  $Z_2$  symmetry is maintained. The top yukawa coupling is obtained by the superpotential

$$W = y_t \mathcal{H}\bar{Q}_R Q_3 \to y_t H_u \frac{1}{\sqrt{2}} \left( \bar{Q}_{R,1} + \bar{Q}_{R,2} \right) Q_3.$$
(2.25)

One linear combination of  $\bar{Q}_{R,1}$  and  $\bar{Q}_{R,2}$  obtains a Dirac mass term paired with U,

$$W = y\bar{Q}_R U(S+\bar{S}) \to yv_S \frac{1}{\sqrt{2}}(\bar{Q}_{R,1} - \bar{Q}_{R,2})U.$$
(2.26)

We identify the massless combination  $(\bar{Q}_{R,1} + \bar{Q}_{R,2})/\sqrt{2} \equiv \bar{u}_3$  as a right-handed top quark. The extra SU(2)<sub>X</sub> charged particle  $\bar{E}$  obtains its mass paired with  $E_{1,2}$  through the coupling,

$$W = \bar{E}(y_{E,1}E_1 + y_{E,2}E_2)S + \bar{E}(y_{E,1}E_1 - y_{E,2}E_2)\bar{S}.$$
(2.27)

Here we assume that  $E_2$  is odd under the  $Z_2$  symmetry  $S \leftrightarrow \overline{S}$ , so that all particles in  $\overline{E}$  and  $E_{1,2}$  obtains their masses.

So far we have assumed that a linear combination of  $\bar{Q}_{R,1}$  and  $\bar{Q}_{R,2}$  obtains a large mass paired with U. It is also possible to identify the linear combination with the right-handed charm quark. In such a model U and  $\bar{u}_2$  are not necessary. The mass of the right-handed scharm is predicted to be as large as that of the right-handed stop. This choice is beneficial for a high mediation scale, as it makes the SU(3)<sub>c</sub> and U(1)<sub>Y</sub> coupling constants relatively smaller, reducing the fine-tuning from the gluino and the bino.

#### 3 Fine-tuning of the electroweak scale

Let us now discuss fine-tuning of the EW scale in the model. We quantify the degree of fine-tuning by introducing the measure [35],

$$\Delta_v \equiv \Delta_f \times \Delta_{v/f},\tag{3.1}$$

where

$$\Delta_{v/f} = \frac{1}{2} \left( \frac{f^2}{v^2} - 2 \right), \tag{3.2}$$

$$\Delta_f = \max_i \left( \left| \frac{\partial \ln f^2}{\partial \ln x_i(\Lambda)} \right|, 1 \right).$$
(3.3)

Here  $f \equiv \sqrt{v^2 + v'^2}$  is the decay constant of the spontaneous SU(4) breaking.  $\Delta_{v/f}$  measures sures the fine-tuning to obtain v < f via explicit soft  $\mathbb{Z}_2$  symmetry breaking.  $\Delta_f$  measures the fine-tuning to obtain the scale f from the soft SUSY breaking which is analogous to the fine-tuning to obtain the electroweak scale from the soft SUSY breaking in the MSSM.  $x_i(\Lambda)$  are the parameters of the theory evaluated at the mediation scale of the SUSY breaking  $\Lambda$ . We include the important seven parameters,  $m_{H_u}^2$ ,  $m_{Q_3}^2$ ,  $m_{\tilde{u}_3}^2$ ,  $M_1^2$ ,  $M_2^2$ ,  $M_3^2$  and  $\mu^2$ . To evaluate  $\Delta_f$  we solve the renormalization group equations (RGEs) of parameters between  $m_{\text{stop}}$  and  $\Lambda$ . We assume that the right-handed charm quark is also embedded in  $\bar{Q}_R$ . Between  $m_{\text{stop}}$  and  $m_X$  we solve MSSM RGEs at the one-loop level appropriately modifying the beta function of  $m_{Q_3}^2$ . At a scale  $m_X$  we perform matching by including the threshold correction (2.11) to  $m_{H_u}^2$  and  $m_{U_3}^2$ . Above  $m_X$  we solve the RGEs (that include the effects of non-MSSM states) at least at the one-loop level. The RGEs of the gauge couplings are solved at the two-loop level, but set, for simplicity,  $\kappa = 0$ .<sup>2</sup> The yukawa couplings other than the top yukawa are neglected.

As clearly seen from eqs. (3.1)–(3.3), for a given value of f there is a lower bound on  $\Delta_v$  of  $\Delta_{v/f}$ . f/v is constrained by the Higgs coupling measurements [38] to be at least 2.3 [39]. The latter value has been obtained neglecting invisible decays of the Higgs to mirror particles, which are generically non-negligible, so in our numerical analysis we use less extremal value of f = 3v. Nevertheless, the tuning is quite independent of this choice (unless f is so large that  $\Delta_{v/f}$  determines  $\Delta_v$ ).

In figure 2 we present contours of  $\Delta_v$  assuming low and high mediation scales of SUSY breaking  $\Lambda$ .<sup>3</sup> Here and hereafter, the stop mass  $m_{\text{stop}}$  and the gluino mass  $M_3$  refer to the values at the TeV scale. For  $\Lambda = 100m_{\text{stop}}$  tuning at the level of 10% can be obtained for the stop masses as large as 3 TeV, as seen from the upper left panel. An important constraint on the parameter space is provided by the Higgs mass measurement [40]. In order to assess the impact of this constraint we compute the Higgs mass following closely the procedure described in ref. [37]. The blue bands show the parameter region with

<sup>&</sup>lt;sup>2</sup>Non-zero  $\kappa$  slightly slows down the running of  $g_X$  but the impact on  $\Delta_v$  and the scale of the Landau pole is negligible.

<sup>&</sup>lt;sup>3</sup>In the figure we shade the parameter region where the Landau pole scale of the gauge coupling  $g_X$  is above  $\Lambda$ . It is also possible that the SUSY breaking is mediated above the Landau pole scale, but we cannot calculate the fine-tuning measure unless we specify the description of the model above the Landau pole scale.





Figure 2. Fine-tuning (red contours) in the model for f = 3v,  $\mu = M_1 = M_2 = 500 \text{ GeV}$ ,  $m_A = 1 \text{ TeV}$  and the soft gluino mass term  $M_3 = 2 \text{ TeV}$  assuming the mediation scale  $\Lambda = 100m_{\text{stop}}$  (upper panels) and  $\Lambda = 10^{16} \text{ GeV}$  (lower panels). In the left panels, the orange contours depict the value of the SU(4) preserving quartic coupling and in the green regions the Landau pole of the SU(2)<sub>X</sub> gauge coupling constant is below  $\Lambda$ . In the upper (lower) left panels,  $\tan \beta = 2.5$  (3) so that the correct Higgs mass  $m_h = 125 \pm 3 \text{ GeV}$  (the blue region) is obtained for stop masses close to 2 TeV for the most interesting range of  $g_X$ . In the right panels, the fine-tuning is shown in the plane  $m_{\text{stop}}$ -tan  $\beta$  for some fixed values of  $g_X$ .  $m_X$  is chosen such that the constraint from EW precision measurements is saturated — see appendix for details.

 $m_h = 125 \pm 3 \text{ GeV}$ , where the error is a theoretical one. It can be seen from the upper right panel of figure 2 that this constraint prefers rather light stop unless  $\tan \beta$  is small enough. Since we are most interested in stop masses that easily avoid current or even potential future LHC constraint we set for the low scale mediation case  $\tan \beta = 2.5$  which implies the stop masses in the range between about 1.5 and 3 TeV. This range narrows to between

1.7 and  $2 \,\mathrm{TeV}$  if one demands tuning better than 10%. Interestingly, tuning is minimised for intermediate values of the stop masses which is a consequence of some cancellation between the threshold correction from X and corrections from stops and gluino to  $m_{H_{\pi}}^2$ . In this region the value of  $|m_{H_u}^2|$  at the mediation scale is somewhat suppressed. For lighter stops (which can be compatible with the Higgs mass constraint for larger tan  $\beta$ ) the tuning is dominated by the threshold correction which implies tuning at the level of few percent. It should be noted that fine-tuning of the EW scale is minimized at some intermediate value of  $g_X$  of about 1.5–2 even though perturbativity constraint allows for  $g_X$  as large as about 2.5. This is because for appropriately large  $g_X$  the tuning is dominated by the threshold correction to  $m_{H_u}^2$  from the new gauge bosons. Since the latter must be rather heavy for large  $g_X$  due to EW precision constraints, the threshold correction dominates for  $g_X \gtrsim 2$  and the tuning gets worse with increasing  $g_X$  in spite of larger SU(4) invariant coupling. In fact, for very large value of  $g_X$  there is essentially no tuning of the EW scale from stops and gluino but the overall tuning is at the level of few percent. In the region of large  $g_X$ , where the threshold correction dominates the fine-tuning, larger values of  $\epsilon$ lead to smaller tuning. On the other hand, for smaller  $g_X$ , when the threshold correction is subdominant, it is preferred to have smaller  $\epsilon$  to suppress corrections from stops and gluino by larger SU(4) invariant coupling.

It is interesting to compare the fine-tuning of the present model to that in the model where an SU(4) invariant coupling originates from a non-decoupling *D*-term of U(1)<sub>X</sub> gauge symmetry proposed in ref. [37]. For the stop mass below about 1 TeV, the U(1)<sub>X</sub> is less tuned with tuning even better than 20%. This is because the threshold correction from the X gauge bosons in the U(1)<sub>X</sub> case is three times smaller than in the case of SU(2)<sub>X</sub>. As the stop mass increases the tuning in the U(1)<sub>X</sub> model gets worse and already for 2 TeV stops the tuning in the SU(2)<sub>X</sub> model becomes better than in the U(1)<sub>X</sub> model due to larger SU(4) invariant coupling which suppresses the correction from stops.

The biggest advantage of the  $SU(2)_X$  model is that RGE running of  $q_X$  is relatively slow so the Landau pole scale, for given  $g_X$ , is much higher than in the  $U(1)_X$  model. For example in the case of  $\Lambda = 10^{16} \,\text{GeV}$  presented in the lower panels of figure 2, values of  $q_X$  up to about 1.2 are possible without the Landau pole below  $\Lambda$ . In the previously proposed SUSY Twin Higgs models it is was not possible to keep perturbativity up to such high scale. We see from figure 2 that for  $\Lambda = 10^{16} \,\text{GeV}$  the fine-tuning better than few % can be obtained for the stop masses as large as 2 TeV. This is obviously worse than in the low-scale mediation case discussed before but for high-scale mediation there are more possible mechanisms of the mediation of the SUSY breaking. The fine-tuning is also much better than in the MSSM with high-scale mediation. This is due to suppression of the corrections from stops and gluino (which dominates tuning for high mediation scales) by the SU(4) invariant coupling but also because a large value of  $g_X$  efficiently drives the top yukawa coupling to smaller values at higher scales. Dependence of fine-tuning on the mediation scale for 2 TeV stops is presented in figure 3. We see that moderate tuning of few percent can be obtained for high mediation scales. For high mediation scales the tuning is dominated by the correction from the gluino so the tuning crucially depends on the gluino mass limits. It was recently emphasised in ref. [6] that one should convert



Figure 3. The same as in figure 2 but in the plane  $\Lambda$ - $g_X$  for  $m_{\text{stop}} = 2 \text{ TeV}$ ,  $\tan \beta = 3$  and  $M_3 = 2 \text{ TeV}$  (left panel) or 2.5 TeV (right panel). For the chosen values of  $m_{\text{stop}}$  and  $\tan \beta$ , the Higgs mass is in agreement with the measured value within theoretical uncertainties in the most of parameter space. For  $g_X \gtrsim 1.5$  the Higgs mass is slightly too big which can be compensated by reducing  $\tan \beta$  by about 10 % which would have negligible impact on fine-tuning.

running soft masses to pole masses when assessing the impact of experimental constraints on naturalness of SUSY models. It was shown that the loop corrections [41] from 2 TeV squarks increase the gluino pole mass by 10% as compared to the soft mass. For heavier 1st/2nd generation of squarks, as experimentally preferred, the correction may be much larger e.g. 20% for 10 TeV squarks. In the left panel of figure 3 we fix the soft gluino mass to 2 TeV which easily satisfies the LHC constraints even for moderate loop corrections from squarks [42, 43]. In such a case, 5% tuning is possible with the mediation scale, being below the Landau pole scale, as high as  $\mathcal{O}(10^{12})$  GeV. For  $M_3 = 2.5$  TeV, presented in the right panel, for which the gluino is definitely outside of the LHC reach [44], mediation scale of order  $\mathcal{O}(10^{10})$  GeV can still allow for better than 5% tuning. Notice also a sharp increase in tuning when the mediation scale approaches the Planck scale. This originates from the fact that  $U(1)_Y$  gauge coupling constant runs rather fast due to many new states carrying hypercharge and eventually enters non-perturbative regime around the Planck scale. In consequence, bino strongly dominates fine-tuning when the mediation scale is close to the Landau pole for  $U(1)_Y$ .

The fine-tuning for high mediation scales is even better if the gluino obtains its mass paired with an adjoint chiral superfield by a supersoft operator, due to the absence of the log-enhanced correction to  $m_{H_u}^2$  [45]. The soft stop mass and the higgs mass are dominantly generated by the threshold correction around the gluino mass,

$$m_{\rm stop}^2 \simeq \frac{1}{16} M_3^2,$$
 (3.4)

$$m_{\rm H_u}^2 \simeq \frac{3y_t(M_3)^2}{4\pi^2} m_{\rm stop} \ln \frac{M_3}{m_{\rm stop}} \,.$$
 (3.5)



Figure 4. The same as in the left panel of figure 3 but with a Dirac gluino with a soft mass  $M_3 = 8 \text{ TeV}$  and  $M_1 = M_2 = 200 \text{ GeV}$ .

In non-Twin models the fine-tuning may be at a few % level even if the stop mass is as large as 2 TeV, which is further improved by the Twin-Higgs mechanism. The contour of  $\Delta_v$ assuming the Dirac gluino is shown in figure 4. For the stop mass of 2 TeV,  $\mathcal{O}(10)$ % tuning is possible even if the mediation scale is as high as  $10^{16}$  GeV. Note that in Dirac gluino models the large log enhancement of the quantum correction to the Higgs mass squared is already absent. Thus the improvement of the fine-tuning by the Twin higgs mechanism simply originates from a large SU(4) invariant coupling. For  $g_X = 1-1.5$ , the improvement is by a factor of 2–4.

In some UV completions of the Dirac gluino, the fine-tuning may be worse and at the O(1)% level [46]. For example in gauge mediated models, a tachyonic soft mass term of the adjoint chiral superfield larger than the Dirac gluino mass is often generated. See ref. [47] for a pedagogical discussion. To prevent the instability of the adjoint field one needs to cancel the tachyonic mass by additional large soft mass or a supersymmetric mass of the adjoint, which leads to fine-tuning. See ref. [48] for a gauge mediated model free from this problem. In gravity mediated model the tachyonic mass is not necessarily larger than the Dirac gluino mass. Our D-term model, together with the Dirac gluino, realizes the natural SUSY even for the gravity mediation.

The wino and the bino masses are also bounded from above by naturalness. The constraint is stronger than that in the MSSM as we add extra  $SU(2)_L$  and/or  $U(1)_Y$  charged fields which makes the corresponding gauge couplings and gaugino masses growing faster with the renormalization scale. Fine-tuning from bino and wino may be very large especially for high mediation scales. In the left panel of figure 5 we fix  $\Lambda = 10^{17} \text{ GeV}$  and present contours of fine-tuning in the plane  $M_1 - M_2$ . We see that bino as light as 700 GeV induces tuning at the level of 1 % for this mediation scale. The tuning from wino is slightly smaller but still 1 TeV wino results in about 1 % tuning. From the comparison



**Figure 5.** Fine-tuning in the plane  $M_1 - M_2$  with a Dirac gluino for  $\Lambda = 10^{17}$  GeV and maximal value of the SU(2)<sub>X</sub> gauge coupling constant,  $g_X^{\text{max}}$ , that do not induce the Landau pole below the mediation scale (left panel). In the right panel, fine-tuning in the plane  $\Lambda - M_2$  is shown for  $M_1 = 100$  GeV and  $g_X = \min(1.5, g_X^{\text{max}})$ . The remaining parameters are the same as in figure 4.

of figures 4 and 5 we see that in order not to increase tuning by more than a factor of two, bino (wino) must be lighter than about 400 (600) GeV. Thus, one generally expects all neutralinos to be light and the LSP to be a mixture of bino, wino and higgsino. Assuming majorana gluino, the impact of wino and bino on the tuning is less pronounce but in order not to increase tuning by more than a factor of two their masses are still expected to be below about 1 TeV, cf. figures 3 and 5. For smaller mediation scales the tuning from bino and wino is milder. The tuning from bino is subdominant unless  $\Lambda \gtrsim 10^{16}$  GeV. In the right panel of figure 5 we present tuning from wino as a function of the mediation scale. We see that even for small mediation scale wino mass should generally be below 1 TeV in order not to dominate tuning. The bounds on the masses is avoided if the wino and the bino also obtain Dirac masses. Interestingly, with an additional SU(2)<sub>L</sub> adjoint paired with wino, the SU(2)<sub>L</sub> gauge coupling constant also blows up around the Planck scale.

In the above analysis we have ignored the contribution to the RGE running of  $m_{H_u}^2$  proportional to the SU(2)<sub>X</sub> gauge coupling constant. As long as the SU(2)<sub>X</sub> gaugino mass is suppressed, one-loop contributions are negligible. At the two loop level, there is a contribution,

$$\frac{\mathrm{d}}{\mathrm{dln}\mu}m_{H_u}^2 \supset \frac{3g_X^4}{256\pi^4}\sum_i m_i^2,\tag{3.6}$$

where  $m_i^2$  is a soft mass squared of a SU(2)<sub>X</sub> fundamental. Although this is a two-loop effect, the largeness of  $m_S^2$  required to obtain a large non-decoupling SU(4) invariant quartic and the largeness of  $g_X$  around the Landau pole scale can make this contribution non-negligible. In the left panel of figure 6, we show the fine-tuning including this two-loop effect



Figure 6. The same as in the left panel of figure 3 but including the effect of  $m_S^2$  under the assumption that it is constant during the RGE flow (left panel) or assuming that  $m_S^2 = 0$  at the mediation scale and large value of  $m_S^2$  at low energy corresponding to  $\epsilon^2 = 0.1$  is obtained due to RG contribution from the soft mass  $m_Z^2$  for  $\kappa = 0.3$  (right panel).

to  $m_{H_u}^2$ , with  $m_S^2$  fixed at the value determined by eq. (2.6), while ignoring contribution from other  $SU(2)_X$  charged fields. The fine-tuning gets worse than the case ignoring the two-loop effect, especially when the mediation scale is close to the Landau pole scale, while it remains the same if the mediation scale is much smaller than the Landau pole scale. We note, however, that the two-loop effect strongly depends on the boundary condition of soft masses at the mediation scale and might be much smaller. For example, if  $m_S^2 = -m_{\bar{E}}^2$  at the boundary, the two-loop effect is suppressed. This special boundary condition should be explained by a UV completion of our model. It is also possible that  $m_S^2$  at the UV boundary is much smaller than around the  $m_X$  scale, and is generated through the RG running. Actually couplings with the fields Z and/or  $E_{1,2}$  can generate a non-zero and positive  $m_S^2$ , if the soft masses of them are negative. For example, in the right panel of figure 6 we show tuning under the assumption that  $m_S^2$  vanishes at the mediation scale and gets renormalized to appropriately large value determined by eq. (2.6) at the EW scale via the interaction (2.1) (by suitable choice of the soft mass for Z) with  $\kappa = 0.3$  at the UV boundary. In this case the impact of  $m_S^2$  on tuning is rather small unless the mediation scale is very close to the Landau pole scale.

Even though naturalness does not require sparticles to be within discovery reach of the LHC (perhaps except for wino if the mediation scale is high enough) it does require that the twin Higgs boson is relatively light. The mass of the twin Higgs boson is well approximated by  $2\sqrt{\lambda}f$  with  $\lambda$  being the SU(4) invariant coupling. Both  $\lambda$  and f are constrained from above by naturalness. For example, demanding better than 10% tuning f must be below about 4.5. This upper bound on f is quite generic for Twin Higgs models unless hard  $\mathbb{Z}_2$  breaking is non-negligible. The upper bound on  $\lambda$  is specific for this model and is set by

the requirement of not too large threshold correction from X gauge bosons. We find that better than 10% tuning requires  $\lambda \leq 0.5$  which, together with the upper bound on f, leads to the upper bound on the twin Higgs boson mass of about 1 TeV. The twin Higgs tends to be lighter for a larger Landau pole scale. For recent studies of the phenomenology of the twin Higgs boson we refer the reader to refs. [36, 39]. It is also noteworthy that in this model MSSM-like Higgs bosons and their mirror counterparts are not required to be light by naturalness because  $H_d$  is not charged under  $SU(2)_X$ .

### 4 Summary

We proposed a new SUSY Twin Higgs model in which an SU(4) invariant quartic term originates from a *D*-term potential of a new SU(2)<sub>X</sub> gauge symmetry. The choice of the non-abelian gauge symmetry, together with a minimal number of flavors charged under SU(2)<sub>X</sub>, makes the running of the new gauge coupling constant rather slow allowing for a large SU(4) invariant quartic term without generating a low-scale Landau pole. The Twin Higgs mechanism, together with the negative contribution from the new gauge coupling to the RG running of the top yukawa coupling, allows for tuning of the EW scale better than 10% for high mediation scales up to  $\mathcal{O}(10^9)$  GeV even for sparticle spectra that may be outside of the ultimate LHC reach. If the gluino obtains a Dirac mass term, tuning of 10% is possible even if the mediation scale is around the Planck scale. The model may be tested at the LHC by searching for a twin Higgs boson whose mass is bounded from above by naturalness and is anti-correlated with the Landau Pole scale. In parts of parameter space with tuning better than 10% the twin Higgs boson is expected to be lighter than about 1 TeV. All electroweakinos are expected to be rather light, with masses in the sub-TeV region, especially if the mediation scale of SUSY breaking is high.

## Acknowledgments

This work has been partially supported by National Science Centre, Poland, under research grant DEC-2014/15/B/ST2/02157, by the Office of High Energy Physics of the U.S. Department of Energy under Contract DE-AC02-05CH11231, and by the National Science Foundation under grant PHY-1316783. MB acknowledges support from the Polish Ministry of Science and Higher Education through its programme Mobility Plus (decision no. 1266/MOB/IV/2015/0). The work of KH was in part performed at the Aspen Center for Physics, which is supported by National Science Foundation grant PHY-1607611.

#### A Electroweak precision measurements

We use the so-called S, T, U parametrization [50] to constrain the parameter space of our model. We follow the method presented in [51], where the observables shown in table 2 are used to constrain S, T, U. We take U = 0 and show the constraint on (S, T) in figure 7.

The Higgs multiplet  $\mathcal{H}$  is charged under  $\mathrm{SU}(2)_X$ . After the electroweak symmetry breaking scale, Z boson mixes with the  $\mathrm{SU}(2)_X$  gauge bosons. The mixing breaks the

| observables                          | value                         | reference |
|--------------------------------------|-------------------------------|-----------|
| $\alpha^{-1}$                        | 137.035999139(31)             | [52]      |
| $G_F(\text{GeV}^{-2})$               | $1.1663787(6) \times 10^{-5}$ | [52]      |
| $m_Z \ ({\rm GeV})$                  | 91.1876(21)                   | [53]      |
| $\Delta \alpha_{ m lep}(m_Z^2)$      | 0.03150                       | [54]      |
| $\Delta \alpha_{\rm had}(m_Z^2)$     | 0.02764(13)                   | [53]      |
| $m_W ~({ m GeV})$                    | 80.385(15)                    | [53]      |
| $m_t ~({\rm GeV})$                   | 173.3(8)                      | [55]      |
| $m_h \; (\text{GeV})$                | 125                           | [40]      |
| $ar{s}_\ell^2$                       | 0.23153(16)                   | [56]      |
| $\Gamma_{Z \to \ell^+ \ell^-}$ (MeV) | 83.984(86)                    | [53]      |
| $\alpha_s(m_Z^2)$                    | 0.1181(11)                    | [53]      |

**Table 2.** Values of observable use to constraint the STU parameters.



Figure 7. The constraint on the S,T parameters.

custodial symmetry and we expect a severe constraint from the electroweak precision measurement. After integrating out the  $SU(2)_X$  gauge bosons, we obtain the effective dimension 6 operator,

$$\mathcal{L}_{\text{eff}} = \frac{g_X^2}{8m_X^2} \left( H_u^{\dagger} D_{\mu} H_u - (D_{\mu} H_u)^{\dagger} H_u \right)^2.$$
(A.1)

This generates a non-zero T parameter,

$$T = \frac{1}{2\alpha} \frac{c_W^2}{g_2^2} \frac{g_X^2 m_Z^2}{m_X^2} \sin^2 \beta \,. \tag{A.2}$$

The dependence on  $\tan \beta$  originates from subdominant  $H_d$  component of the Higgs. S and U parameters are negligibly small. Comparing this result with figure 7, we obtain the constraint,

$$m_X/g_X > 4.1 \text{ TeV}\sin^2\beta. \tag{A.3}$$

The mixing between the SM-like Higgs and the mirror Higgs also contributes to S and T parameters,

$$S = \frac{1}{12\pi} s_{\gamma}^2 \ln \frac{m_{h'}^2}{m_h^2},\tag{A.4}$$

$$T = -\frac{3}{16\pi c_w^2} s_\gamma^2 \ln \frac{m_{h'}^2}{m_h^2}.$$
 (A.5)

where  $\gamma$  is the mixing angle between the SM Higgs and the mirror Higgs. Note that the sign of the T parameter is negative, and hence this contribution relaxes the constraint on the  $SU(2)_X$  symmetry breaking scale. In section 3, we use the constraint including the Higgs mixing to determine the magnitude of the threshold correction around the  $SU(2)_X$  symmetry breaking scale.

**Open Access.** This article is distributed under the terms of the Creative Commons Attribution License (CC-BY 4.0), which permits any use, distribution and reproduction in any medium, provided the original author(s) and source are credited.

#### References

- L. Maiani, All You Need to Know about the Higgs Boson, in proceedings of Summer School on Particle Physics, Paris, France (1979) [INSPIRE].
- M.J.G. Veltman, The Infrared-Ultraviolet Connection, Acta Phys. Polon. B 12 (1981) 437 [INSPIRE].
- [3] E. Witten, Dynamical Breaking of Supersymmetry, Nucl. Phys. B 188 (1981) 513 [INSPIRE].
- [4] R.K. Kaul, Gauge Hierarchy in a Supersymmetric Model, Phys. Lett. B 109 (1982) 19 [INSPIRE].
- [5] M.R. Buckley, D. Feld, S. Macaluso, A. Monteux and D. Shih, Cornering Natural SUSY at LHC Run II and Beyond, JHEP 08 (2017) 115 [arXiv:1610.08059] [INSPIRE].
- [6] M.R. Buckley, A. Monteux and D. Shih, Precision Corrections to Fine Tuning in SUSY, JHEP 06 (2017) 103 [arXiv:1611.05873] [INSPIRE].
- S. Dimopoulos, K. Howe and J. March-Russell, Maximally Natural Supersymmetry, Phys. Rev. Lett. 113 (2014) 111802 [arXiv:1404.7554] [INSPIRE].
- [8] T. Cohen, J. Kearney and M. Luty, Natural Supersymmetry without Light Higgsinos, Phys. Rev. D 91 (2015) 075004 [arXiv:1501.01962] [INSPIRE].
- [9] S.P. Martin, Nonstandard supersymmetry breaking and Dirac gaugino masses without supersoftness, Phys. Rev. D 92 (2015) 035004 [arXiv:1506.02105] [INSPIRE].
- [10] I. Garcia Garcia, K. Howe and J. March-Russell, Natural Scherk-Schwarz Theories of the Weak Scale, JHEP 12 (2015) 005 [arXiv:1510.07045] [INSPIRE].

- [11] Z. Chacko, H.-S. Goh and R. Harnik, The Twin Higgs: Natural electroweak breaking from mirror symmetry, Phys. Rev. Lett. 96 (2006) 231802 [hep-ph/0506256] [INSPIRE].
- [12] Z. Chacko, Y. Nomura, M. Papucci and G. Perez, Natural little hierarchy from a partially goldstone twin Higgs, JHEP 01 (2006) 126 [hep-ph/0510273] [INSPIRE].
- [13] Z. Chacko, H.-S. Goh and R. Harnik, A Twin Higgs model from left-right symmetry, JHEP 01 (2006) 108 [hep-ph/0512088] [INSPIRE].
- [14] A. Falkowski, S. Pokorski and M. Schmaltz, Twin SUSY, Phys. Rev. D 74 (2006) 035003
   [hep-ph/0604066] [INSPIRE].
- [15] S. Chang, L.J. Hall and N. Weiner, A Supersymmetric twin Higgs, Phys. Rev. D 75 (2007) 035009 [hep-ph/0604076] [INSPIRE].
- [16] P. Batra and Z. Chacko, A Composite Twin Higgs Model, Phys. Rev. D 79 (2009) 095012 [arXiv:0811.0394] [INSPIRE].
- [17] M. Geller and O. Telem, *Holographic Twin Higgs Model*, *Phys. Rev. Lett.* **114** (2015) 191801
   [arXiv:1411.2974] [INSPIRE].
- [18] R. Barbieri, D. Greco, R. Rattazzi and A. Wulzer, The Composite Twin Higgs scenario, JHEP 08 (2015) 161 [arXiv:1501.07803] [INSPIRE].
- [19] M. Low, A. Tesi and L.-T. Wang, Twin Higgs mechanism and a composite Higgs boson, Phys. Rev. D 91 (2015) 095012 [arXiv:1501.07890] [INSPIRE].
- [20] H.-C. Cheng, S. Jung, E. Salvioni and Y. Tsai, Exotic Quarks in Twin Higgs Models, JHEP 03 (2016) 074 [arXiv:1512.02647] [INSPIRE].
- [21] C. Csáki, M. Geller, O. Telem and A. Weiler, The Flavor of the Composite Twin Higgs, JHEP 09 (2016) 146 [arXiv:1512.03427] [INSPIRE].
- [22] H.-C. Cheng, E. Salvioni and Y. Tsai, Exotic electroweak signals in the twin Higgs model, Phys. Rev. D 95 (2017) 115035 [arXiv:1612.03176] [INSPIRE].
- [23] R. Contino, D. Greco, R. Mahbubani, R. Rattazzi and R. Torre, Precision Tests and Fine Tuning in Twin Higgs Models, arXiv:1702.00797 [INSPIRE].
- [24] R. Barbieri, L.J. Hall and K. Harigaya, Minimal Mirror Twin Higgs, JHEP 11 (2016) 172 [arXiv:1609.05589] [INSPIRE].
- [25] Z. Chacko, N. Craig, P.J. Fox and R. Harnik, Cosmology in Mirror Twin Higgs and Neutrino Masses, JHEP 07 (2017) 023 [arXiv:1611.07975] [INSPIRE].
- [26] N. Craig, S. Koren and T. Trott, Cosmological Signals of a Mirror Twin Higgs, JHEP 05 (2017) 038 [arXiv:1611.07977] [INSPIRE].
- [27] I. Garcia Garcia, R. Lasenby and J. March-Russell, Twin Higgs WIMP Dark Matter, Phys. Rev. D 92 (2015) 055034 [arXiv:1505.07109] [INSPIRE].
- [28] I. Garcia Garcia, R. Lasenby and J. March-Russell, Twin Higgs Asymmetric Dark Matter, Phys. Rev. Lett. 115 (2015) 121801 [arXiv:1505.07410] [INSPIRE].
- [29] N. Craig and A. Katz, The Fraternal WIMP Miracle, JCAP 10 (2015) 054 [arXiv:1505.07113] [INSPIRE].
- [30] M. Farina, Asymmetric Twin Dark Matter, JCAP 11 (2015) 017 [arXiv:1506.03520] [INSPIRE].

- [31] M. Freytsis, S. Knapen, D.J. Robinson and Y. Tsai, Gamma-rays from Dark Showers with Twin Higgs Models, JHEP 05 (2016) 018 [arXiv:1601.07556] [INSPIRE].
- [32] M. Farina, A. Monteux and C.S. Shin, Twin mechanism for baryon and dark matter asymmetries, Phys. Rev. D 94 (2016) 035017 [arXiv:1604.08211] [INSPIRE].
- [33] V. Prilepina and Y. Tsai, Reconciling Large And Small-Scale Structure In Twin Higgs Models, JHEP 09 (2017) 033 [arXiv:1611.05879] [INSPIRE].
- [34] R. Barbieri, L.J. Hall and K. Harigaya, Effective Theory of Flavor for Minimal Mirror Twin Higgs, arXiv:1706.05548 [INSPIRE].
- [35] N. Craig and K. Howe, Doubling down on naturalness with a supersymmetric twin Higgs, JHEP 03 (2014) 140 [arXiv:1312.1341] [INSPIRE].
- [36] A. Katz, A. Mariotti, S. Pokorski, D. Redigolo and R. Ziegler, SUSY Meets Her Twin, JHEP 01 (2017) 142 [arXiv:1611.08615] [INSPIRE].
- [37] M. Badziak and K. Harigaya, Supersymmetric D-term Twin Higgs, JHEP 06 (2017) 065 [arXiv:1703.02122] [INSPIRE].
- [38] ATLAS, CMS collaborations, Measurements of the Higgs boson production and decay rates and constraints on its couplings from a combined ATLAS and CMS analysis of the LHC pp collision data at  $\sqrt{s} = 7$  and 8 TeV, JHEP **08** (2016) 045 [arXiv:1606.02266] [INSPIRE].
- [39] D. Buttazzo, F. Sala and A. Tesi, Singlet-like Higgs bosons at present and future colliders, JHEP 11 (2015) 158 [arXiv:1505.05488] [INSPIRE].
- [40] ATLAS, CMS collaborations, Combined Measurement of the Higgs Boson Mass in pp Collisions at √s = 7 and 8 TeV with the ATLAS and CMS Experiments, Phys. Rev. Lett. 114 (2015) 191803 [arXiv:1503.07589] [INSPIRE].
- [41] D.M. Pierce, J.A. Bagger, K.T. Matchev and R.-j. Zhang, Precision corrections in the minimal supersymmetric standard model, Nucl. Phys. B 491 (1997) 3 [hep-ph/9606211]
   [INSPIRE].
- [42] CMS collaboration, Search for supersymmetry in multijet events with missing transverse momentum in proton-proton collisions at 13 TeV, Phys. Rev. D 96 (2017) 032003
   [arXiv:1704.07781] [INSPIRE].
- [43] ATLAS collaboration, Search for production of supersymmetric particles in final states with missing transverse momentum and multiple b-jets at  $\sqrt{s} = 13$  TeV proton-proton collisions with the ATLAS detector, ATLAS-CONF-2017-021 (2017).
- [44] J.S. Kim, K. Rolbiecki, R. Ruiz, J. Tattersall and T. Weber, Prospects for natural SUSY, Phys. Rev. D 94 (2016) 095013 [arXiv:1606.06738] [INSPIRE].
- [45] P.J. Fox, A.E. Nelson and N. Weiner, Dirac gaugino masses and supersoft supersymmetry breaking, JHEP 08 (2002) 035 [hep-ph/0206096] [INSPIRE].
- [46] A. Arvanitaki, M. Baryakhtar, X. Huang, K. van Tilburg and G. Villadoro, *The Last Vestiges of Naturalness*, JHEP 03 (2014) 022 [arXiv:1309.3568] [INSPIRE].
- [47] C. Csáki, J. Goodman, R. Pavesi and Y. Shirman, The  $m_D b_M$  problem of Dirac gauginos and its solutions, Phys. Rev. D 89 (2014) 055005 [arXiv:1310.4504] [INSPIRE].
- [48] D.S.M. Alves, J. Galloway, M. McCullough and N. Weiner, Goldstone Gauginos, Phys. Rev. Lett. 115 (2015) 161801 [arXiv:1502.03819] [INSPIRE].

- [49] R. Barbieri, T. Gregoire and L.J. Hall, Mirror world at the large hadron collider, hep-ph/0509242 [INSPIRE].
- [50] M.E. Peskin and T. Takeuchi, Estimation of oblique electroweak corrections, Phys. Rev. D 46 (1992) 381 [INSPIRE].
- [51] M.E. Peskin and J.D. Wells, How can a heavy Higgs boson be consistent with the precision electroweak measurements?, Phys. Rev. D 64 (2001) 093003 [hep-ph/0101342] [INSPIRE].
- [52] P.J. Mohr, D.B. Newell and B.N. Taylor, CODATA Recommended Values of the Fundamental Physical Constants: 2014, Rev. Mod. Phys. 88 (2016) 035009 [arXiv:1507.07956] [INSPIRE].
- [53] PARTICLE DATA GROUP collaboration, C. Patrignani et al., *Review of Particle Physics*, *Chin. Phys.* C 40 (2016) 100001 [INSPIRE].
- [54] M. Steinhauser, Leptonic contribution to the effective electromagnetic coupling constant up to three loops, Phys. Lett. B 429 (1998) 158 [hep-ph/9803313] [INSPIRE].
- [55] ATLAS, CDF, CMS, D0 collaborations, First combination of Tevatron and LHC measurements of the top-quark mass, arXiv:1403.4427 [INSPIRE].
- [56] ALEPH, DELPHI, L3, OPAL, SLD collaborations, LEP Electroweak Working Group, SLD Electroweak Group, SLD Heavy Flavour Group, S. Schael et al., Precision electroweak measurements on the Z resonance, Phys. Rept. 427 (2006) 257 [hep-ex/0509008] [INSPIRE].